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2D Velocity and Temperature Measurements in High Speed Flows Based on Spectrally Resolved Rayleigh Scattering

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2D VELOCITY AND TEMPERATURE MEASUREMENTS IN HIGH SPEED FLOWS BASED ON SPECTRALLY RESOLVED RAYLEIGH SCATTERING

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ABSTRACT. The use of molecular Rayleigh scattering for measurements of gas velocity and temperature is evaluated. Molecular scattering avoids problems associated with the seeding required by conventional laser anemometry and particle image velocimetry. The technique considered herein is based on the measurement of the spectrum of the scattered light. Planar imaging of Rayleigh scattering using a laser light sheet is evaluated for conditions at 30km altitude (typical hypersonic flow conditions). The Cramer-Rao lower bounds for velocity and temperature measurement uncertainties are calculated for an ideal optical spectrum analyzer and for a planar mirror Fabry-Perot interferometer used in a static, imaging mode. With this technique, a single image of the Rayleigh scattered light from clean flows can be analyzed to obtain temperature and one component of velocity. Experimental results are presented for planar velocity measurements in a Mach 1.3 air jet.

1. Introduction

Laser diagnostics based on molecular scattering are receiving increased attention for use in aerospace test facilities. These methods can provide information directly related to the parameters of the molecular state of the gas including temperature, density, and bulk velocity. The elimination of the need for seeding is an important advantage of molecular scattering methods compared to conventional laser anemometry. Of the various molecular scattering techniques, Rayleigh scattering is the simplest and least expensive to implement. Rayleigh scattering has chiefly been used to obtain gas density by measuring the total Rayleigh scattered power. More information, however, is available from the spectrum of the Rayleigh scattered light.

The Rayleigh scattering spectrum is directly related to the molecular velocity distribution function, which is a function of the thermodynamic parameters (temperature, density, pressure), composition, mean velocity, and turbulence parameters. Light scattered at a single scattering angle provides information related to one component of the molecular velocities. Thus one component of the mean velocity can be determined (from the shift of the spectral peak). The width of the spectrum is related to the gas temperature and turbulence. For low turbulence, gas temperature can be obtained; or, for a known temperature, turbulence intensity can be obtained.

Extraction of the parameters of interest (temperature and velocity) from the spectrum can be accomplished using a sharp cutoff atomic or molecular absorption

filter [Shimizu et al.(1983); Miles et al.(1991)], or with a high resolution spectrometer, such as a Fabry-Perot or Michelson interferometer. Point measurements of Rayleigh scattering spectra with scanning Fabry-Perot interferometers have been used to measure temperature in a hydrogen-air flame [Cattolica et al.(1976); Pitz et al.(1976)] and velocity and temperature in the exhaust of a small hydrogen-oxygen rocket [Seasholtz et al.(1991)]. Rayleigh scattering has been proposed for atmospheric temperature measurements [Schwiesow and Lading(1981); Rees et al.(1984)]. The application of single pass and multi pass scanning Fabry-Perot interferometers for high speed velocity point measurements has been previously analyzed [Seasholtz (1991)]. Imaging velocity measurements of seeded flows using a Michelson interferometer have also been reported by Seiler and Jäger (1983). In this paper we examine an imaging technique to measure velocity and temperature using a planar mirror Fabry-Perot. The technique is an extension of 1D temperatures and number density measurements in a heated air jet [Lock et al.(1992)] and is similar to Fabry-Perot imaging techniques for measurement of instantaneous velocity of solid surfaces [Gidon and Behar (1986)], temperature and velocity in the upper atmosphere [Sivjee et al.(1980)], and solar corona measurements [Chandrasekhar et al.(1981)].

We first analyze a hypothetical planar imaging experiment where the measurement region is illuminated with a laser light sheet, and the Rayleigh scattered light is detected with a 2D array detector. It is assumed that the medium under study is a single component gas with no turbulence and contains no particles. The Cramer-Rao lower bounds for velocity and temperature measurement errors are calculated. We first calculate the error bounds for an ideal spectrum analyzer, which gives a basis with which to compare the performance of spectral measurements made with real instruments. The error bounds are then calculated for a planar mirror Fabry-Perot interferometer used in the static imaging mode. A specific example typical of hypersonic flow environments is discussed.

The use of 2D Rayleigh imaging for velocity measurements is demonstrated with an experiment designed to measure the axial velocity component in a Mach 1.3 supersonic free jet.

2. Theory

In this section we obtain theoretical lower bounds for uncertainty in velocity and temperature measurements for two cases where the scattered light is detected with a 2D array detector. The first case assumes that an ideal instrument is used to measure the spectrum. The second case assumes that the spectrum is measured with a Fabry-Perot interferometer operated in a static, imaging mode (this is the technique described in the Experiment section). Only errors resulting from the Poisson statistics of the detected light are considered. Thus these results represent lower bounds for measurement errors.

2.1. 2D RAYLEIGH IMAGING DIAGNOSTIC

Consider a planar imaging experiment where a laser sheet uniformly illuminates a region in the object plane corresponding to the image of an array detector having N_r rows (measured perpendicular to the beam propagation direction) and N_c columns of pixels. Let the detector pixel size be $L_x \times L_x$ (square pixels) and let the system have a magnification M (i.e., the measurement plane has dimensions $(N_c L_x \times N_r L_x)/M$).

The expected number of detected photons per pixel can be written [Eckbreth(1988)]

$$\langle \text{NPP} \rangle = \frac{E_1}{N_r} \frac{\epsilon n L_x \lambda_o \Omega}{h c M} \left[\frac{d\sigma}{d\Omega} \right] \quad (1)$$

where E_1 is the laser energy, λ_o is the laser wavelength, $(d\sigma/d\Omega)$ is the differential Rayleigh scattering cross section, n is the molecular number density, Ω is the solid angle of the collected light, h is Planck's constant, c is the velocity of light, and ϵ is the overall detection efficiency (including detector quantum efficiency and system losses).

As an example, we consider air at 30km altitude ($T = 230\text{K}$, $p = 0.0118\text{atm}$, $n = 3.83 \times 10^{23} \text{ m}^{-3}$), $\lambda_o = 532\text{nm}$, $E_1 = 1\text{J}$, $L_x = 23\mu\text{m}$, $N_r = 384$, $M\Omega = 0.05$ ($f/4$ at $M=1$), $\epsilon = 0.05$, and $(d\sigma/d\Omega) = 6.1 \times 10^{-32} \text{ m}^2/\text{sr}$. For this case, equation 1 gives $\langle \text{NPP} \rangle = 9.4$ detected photons per pixel.

2.2. SPECTRUM OF RAYLEIGH SCATTERED LIGHT

For a single component, low density gas with a Maxwellian velocity distribution, the normalized spectrum of the scattered light is [Fabelinskii (1968)]

$$S(f) df = \frac{2\sqrt{\pi}}{aK} e^{-[2\pi(f-f_o) - \mathbf{K} \cdot \mathbf{V}]^2 / a^2 K^2} df \quad (2)$$

where $a = (2\kappa T/m)^{1/2}$ (m = molecular mass; κ = Boltzman's constant). The spectral peak is shifted from the laser frequency $f_o = c/\lambda_o$ by $\mathbf{K} \cdot \mathbf{V}/2\pi$, where \mathbf{V} is the mean velocity of the gas, and $\mathbf{K} = \mathbf{k}_s - \mathbf{k}_o$ (\mathbf{k}_s and \mathbf{k}_o being the wave vectors of the scattered and incident light). The magnitude of \mathbf{K} is a function of the scattering angle θ_s and the wavelength λ

$$K = |\mathbf{K}| = (4\pi/\lambda) \sin(\theta_s/2) \quad (3)$$

The observed Rayleigh scattering spectrum can be expressed as a set of counts

$$\langle n_j \rangle = G_R S(f_j) \Delta f \quad (4)$$

which are the expected number of counts in frequency interval f_j to $f_j + \Delta f$; G_R is the expected total number of counts. The spectrum (eq. 4) is thus a function of a set of three unknown parameters that can be expressed as the vector

$$\alpha = [G_R, V_K, T] \quad (5)$$

where V_K is the velocity component along \mathbf{K} (i.e., $V_K = \mathbf{K} \cdot \mathbf{V}/K$).

2.3. CRAMER-RAO LOWER BOUND

The Cramer-Rao lower bound [Whalen(1971)] for the variance of the estimate of the parameter α_i is given by

$$V(\hat{\alpha}_i) = [\Gamma^{-1}]_{ii} \quad (6)$$

where Γ is the Fisher information matrix given (for Poisson statistics) by [Lading and Jensen(1980)]

$$\Gamma_{ij} = \sum_q \frac{1}{\langle n_q \rangle} \frac{\partial \langle n_q \rangle}{\partial \alpha_i} \frac{\partial \langle n_q \rangle}{\partial \alpha_j} \quad (7)$$

If the parameters are uncorrelated, the estimate of the variance of α_i is simply the inverse of the i th diagonal element of the Fisher matrix elements given by

$$V(\hat{\alpha}_i) = \frac{1}{\Gamma_{ii}} \quad (8)$$

2.3.1. Ideal Spectrum Analyzer. If we assume that the parameters are uncorrelated, we can easily evaluate the Fisher matrix elements and use equation 8 to get the Cramer-Rao bounds. Expressed as standard deviations $\sigma(\alpha_i) = [V(\alpha_i)]^{1/2}$, the lower bounds for measurement uncertainties for velocity and temperature are

$$\sigma(V_K) = \frac{a}{(2G_R)^{1/2}}, \quad \sigma(T) = \frac{2T}{(2G_R)^{1/2}} \quad (9)$$

We now consider a subregion where it is assumed the parameters do not vary. The total number of photoelectron counts due to Rayleigh scattering detected from this subregion is then $G_R = \langle NPP \rangle \times \text{number of pixels}$. The size of the subregion determines the spatial resolution of the measurements. For example, assuming a 20×20 array of pixels (which corresponds to $0.5\text{mm} \times 0.5\text{mm}$) and $\langle NPP \rangle = 9.4$ (as given in the above example), we have $G_R = 3760$ counts, $T = 230\text{K}$, and $a = 370\text{m/s}$. Thus the uncertainties in velocity and temperature are $\sigma(V_K) = 4\text{m/s}$ and $\sigma(T) = 5\text{K}$. Note that the velocity uncertainty is proportional to the square root of temperature.

2.3.2. Fabry-Perot Interferometer. The above error estimates were based on the use of an ideal instrument to measure the spectrum. In this section we perform the error estimation based on measurement of the intensity distribution of light imaged through a planar mirror Fabry-Perot interferometer. The expected number of detected photoelectrons $\langle NDP \rangle$ for the q th pixel is

$$\langle NDP_q \rangle = \langle NPP_q \rangle \iint S(f) I_t(f, \theta_r) df dA_{\text{pixel}} \quad (10)$$

where the Fabry-Perot transmission function [Vaughan(1989)] is

$$I_t(\psi) = \frac{1}{1 + F \sin^2(\psi/2)} \quad (11)$$

with ψ being the phase delay of the light for one pass through the interferometer given by

$$\psi = \frac{4\pi\mu df \cos\theta_r}{c} \quad (12)$$

and $F = (2N_R/\pi)^2$, where N_R is the reflective finesse. In this equation, f is the frequency of the light, μ is the refractive index of the medium in the Fabry-Perot cavity (taken as unity here), d is the Fabry-Perot mirror spacing, and θ_r is the angle between the ray and the optic axis. For simplicity, we neglected spectral broadening due to the finite size of the aperture of the collection optics.

We now select, for this example, a particular Fabry-Perot configuration (mirror spacing $d = 15\text{mm}$; finesse $N_R = 20$) and numerically evaluate the Fisher matrix elements

$$\Gamma_{ij} = \sum_q \frac{1}{\langle NDP_q \rangle} \frac{\partial \langle NDP_q \rangle}{\partial \alpha_i} \frac{\partial \langle NDP_q \rangle}{\partial \alpha_j} \quad (13)$$

TABLE 1. Lower bounds for uncertainty in velocity and temperature measurements.

Fringe number	Fringe radius	$\sigma(V_K)$, m/s	$\sigma(T)$, K
1	38 pixels	14	44
2	69	19	49
3	90	22	61
4	106	24	71
5	121	25	68
11	185	20	81
19	286	24	83
ideal		4	5

where the sum is over the same 20×20 pixel subregion used above. The results of this calculation are given in Table 1 for several locations of the subregion (denoted by fringe radii). The velocity uncertainty is about 20m/s (5 times that achievable with ideal spectrum analyzer) and the temperature uncertainty is about 40-80K (10 times that of ideal). In practice, the assumption of constant finesse will not be valid; vignetting reduces the effective number of reflections in the interferometer cavity, thereby decreasing the finesse at larger fringe radii. This means that the finesse must be measured, as is done in the experiment described later in this paper.

2.4. OTHER CONSIDERATIONS

The example presented above considered an idealized measurement situation. There are always other factors that degrade measurement accuracy. For instance, if the gas is made up of an unknown mixture of component gases, the spectrum cannot be uniquely related to the temperature. If the flow is highly turbulent, and the measurement is obtained over a duration much greater than the turbulence time scale, the spectrum is broadened with both temperature and turbulence contributing to the spectral width. Thus, if one of these is known, the other can be determined from the spectral width, but both cannot be simultaneously determined. The Gaussian spectrum used in the example occurs only when the molecular mean-free-path is greater than the wavelength $\Lambda = 2\pi/K$ involved in the scattering process (i.e., $y \equiv \Lambda/\text{mfp} \ll 1$). For values of $y > 1$, collective effects in the molecular velocities become important and must be taken into account in evaluating the spectrum shape. For $y \gg 1$, relatively simple continuum models can be used [Clark (1975)]; the spectrum in this case splits into three peaks and is referred to as Brillouin scattering. For the transition region, $y \approx 1$, the spectral calculations are more involved and kinetic theory models, such as the Tenti S6 model [Tenti et al.(1974)] must be used. In general, ambient conditions fall into the transition region of $y \approx 1$, while the lower densities typical of hypersonic flows have $y \ll 1$.

Particles in the flow will strongly scatter. In some cases where the particle number density is low, the images can be processed to remove the obvious particles. If the particle number density is so large that the particles images overlap and

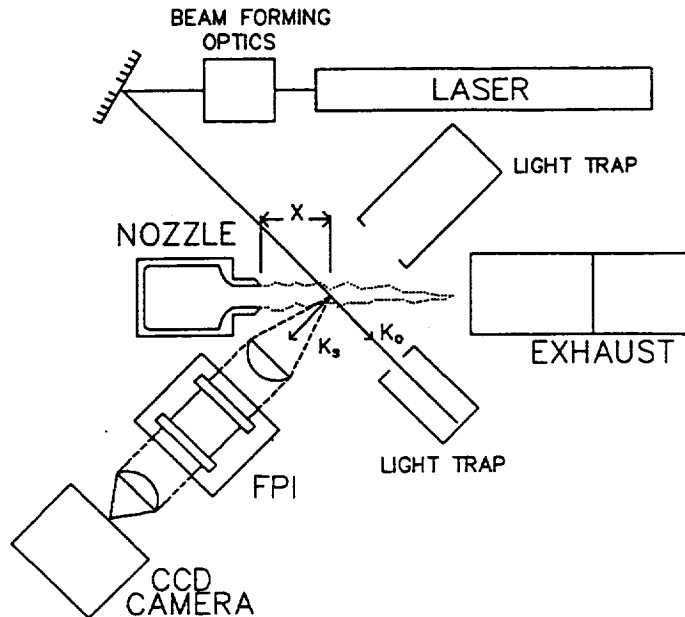


Figure 1. Experimental setup for measurement of velocity in Mach 1.3 free jet using imaging Fabry-Perot interferometer.

cannot be removed, the spectral shape is changed from that of pure molecular scattering and temperature measurements cannot be obtained. It is important to note, however, that particles do not prevent the measurement of velocity (provided there is negligible particle lag) since the spectral peak due to particles is at the same frequency as the peak of the molecularly scattered light.

3. Experiment to Map Velocity in Supersonic Free Jet

3.1. EXPERIMENTAL SETUP

To demonstrate the use of Rayleigh scattering for measuring velocity, we conducted the experiment shown in figure 1. The Mach 1.3 supersonic free jet (9.9mm exit diameter) was mounted on a linear translation stage so it could be moved along its axis. The beam from an argon-ion laser(2.5W, 488nm) was formed into a sheet (about 10mm high by 0.2mm thick) using a combination of cylindrical and spherical lenses. The light sheet crossed the flow at 45°. The receiving optics were aligned at 45° to the flow direction and 90° to the propagation direction of the laser beam. Thus the K vector was aligned with the flow, and the axial velocity component was measured. A 250mm focal length $f/3.6$ lens was used to collect and collimate the scattered light. The collimated light was passed through a planar mirror Fabry-Perot interferometer (70mm diameter aperture and 14.96mm mirror spacing) and focused onto the CCD array with a 225mm focal length lens. The CCD array had 384 rows by 576 columns of $23\mu\text{m} \times 23\mu\text{m}$ square pixels (giving a field size in the flow of $9.8\text{mm} \times 14.7\text{mm}$) and was cooled to about -50°C to reduce dark current. The CCD pixel data was digitized with 12 bit resolution and transferred to an 80386 PC via an IEEE-488 interface for storage and analysis.

The initial step in the measurement process was to measure the instrument function $[I_t(\psi), \text{ for } f=f_0]$. A diffusely scattering target was placed in the field of view and illuminated with laser light scattered from a rotating diffuse target (the

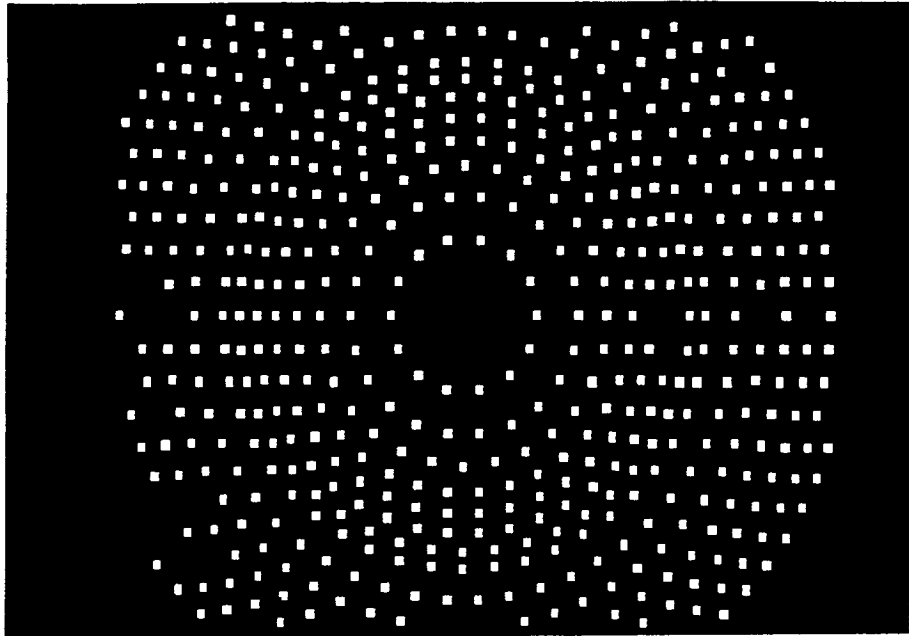


Figure 2. Pattern of subregions used for data analysis (5×5 pixel subregions shown; actual size of regions used was 20×20 pixels).

rotating target was used to reduce speckle in the recorded image. The Rayleigh scattered light from the flow was then recorded. Exposure times were 0.1s for the instrument function and 1.0s for the Rayleigh scattering.

Although the air supplied to the nozzle was filtered, it was not feasible to supply filtered air to the air surrounding the jet. Particles in the ambient air were thus mixed into the flow, and temperature measurements were not realizable (even though the technique described by Kay and Sadler (1991) was used to reduce the effect of particles).

3.2. DATA REDUCTION

The data reduction was done in two steps. First, the instrument function image was analyzed to determine the center of the circular fringe pattern, the finesse, and the phase corresponding to the laser frequency. Because the effective finesse and phase varied over the image plane, it was necessary to determine them as a function of location in the image (typically, the finesse varied from about 20 at the center to about 4 at the edges of the image). About 500 20×20 pixel subregions were used. An integral number were positioned around the each of the inner fringes and around every other outer fringe (so the subregions did not overlap). A typical pattern of subregions is shown in figure 2. A 4th order, 2D polynomial (15 parameters) was then fit to these values of phase and finesse. The coefficients for the polynomial fits were saved for the second step (analyzing the image of the Rayleigh scattered light).

In the second step, the image of the Rayleigh scattered light in each subregion was fit to a model function given by equation 10 using the S6 model of Tenti et al.(1974) with the gas properties of air. The velocity data obtained at each subregion were fit to a 2D 4th order polynomial, which was used to generate contour plots of constant axial velocity.

3.3. RESULTS

Figure 3 shows constant axial velocity contours for 5 axial locations along the jet axis ($X=20,30,50,70$, and 90 mm). Recall that these planes are oriented at 45° to the jet axis. At the locations closer to the nozzle exit, the maximum velocity is in good agreement with the 423m/s velocity calculated from the measured total pressure and temperature (assuming isentropic flow). As the measurement plane moves downstream, the maximum velocity decreases as the core flow mixes with ambient air. The contours are elliptically shaped because of the orientation of the measurement plane. We also can see that the center of the jet is shifting to the right as it moves downstream. This is probably caused by a slight misalignment of the jet axis.

The uncertainty in the velocity estimates due to photon arrival rate statistics is quite small, less than 10m/s . Errors due other sources most certainly outweigh these statistical errors. One potential error source is the drift in the Fabry-Perot mirror spacing between recording the instrument function and the Rayleigh data. The r.m.s. deviation of the contours from the individual subregions data was about $10\text{--}15\text{m/s}$.

The data reduction requires rather lengthy processing, chiefly because of the numerical integration in equation 10. The time per iteration for each subregion is about 15s using a $80386/33\text{MHz}/\text{Weitek } 3167$. Thus an image with 500 subregions and an average of 5 iterations, requires about 10hr to process. No attempt was made to optimize the code at this time, so it should be possible to substantially reduce the data reduction time.

4. Concluding Remarks

The results of this work show the feasibility of using measurements of the spectrum of molecular Rayleigh scattering as a diagnostic suitable for many high speed flow experiments. The technique is capable, under certain restrictions, of mapping gas density, velocity, and temperature.

The experimental work reported here used a CW argon-ion laser, which results in time-averaged measurements. Use of a pulsed laser and gated detection would offer several advantages. Measurements based on a image from a single laser pulse would give 2D maps of the gas parameters at an instant of time. The effect of background light, such as flame luminosity, could be greatly reduced. And, particle images, if not too dense, should be easier to identify and remove from the flow field image.

A large number density of particles in the flow generally will prevent temperature and density measurements. It is important to note, however, that velocity can be obtained with or without particles, provided the particles follow the flow. Care should be taken to prevent spurious laser light from reaching the receiving optics. Even though this light can be accounted for in the data reduction, it will degrade the signal-to-noise ratio and measurement accuracy. This problem is particularly severe in internal flows where it is very difficult to eliminate all spuriously scattered light. The technique described by Miles et al. (1991) for blocking light at the laser frequency should be useful for high velocity flows where the frequency of the Rayleigh scattered light is shifted well away from the laser frequency.

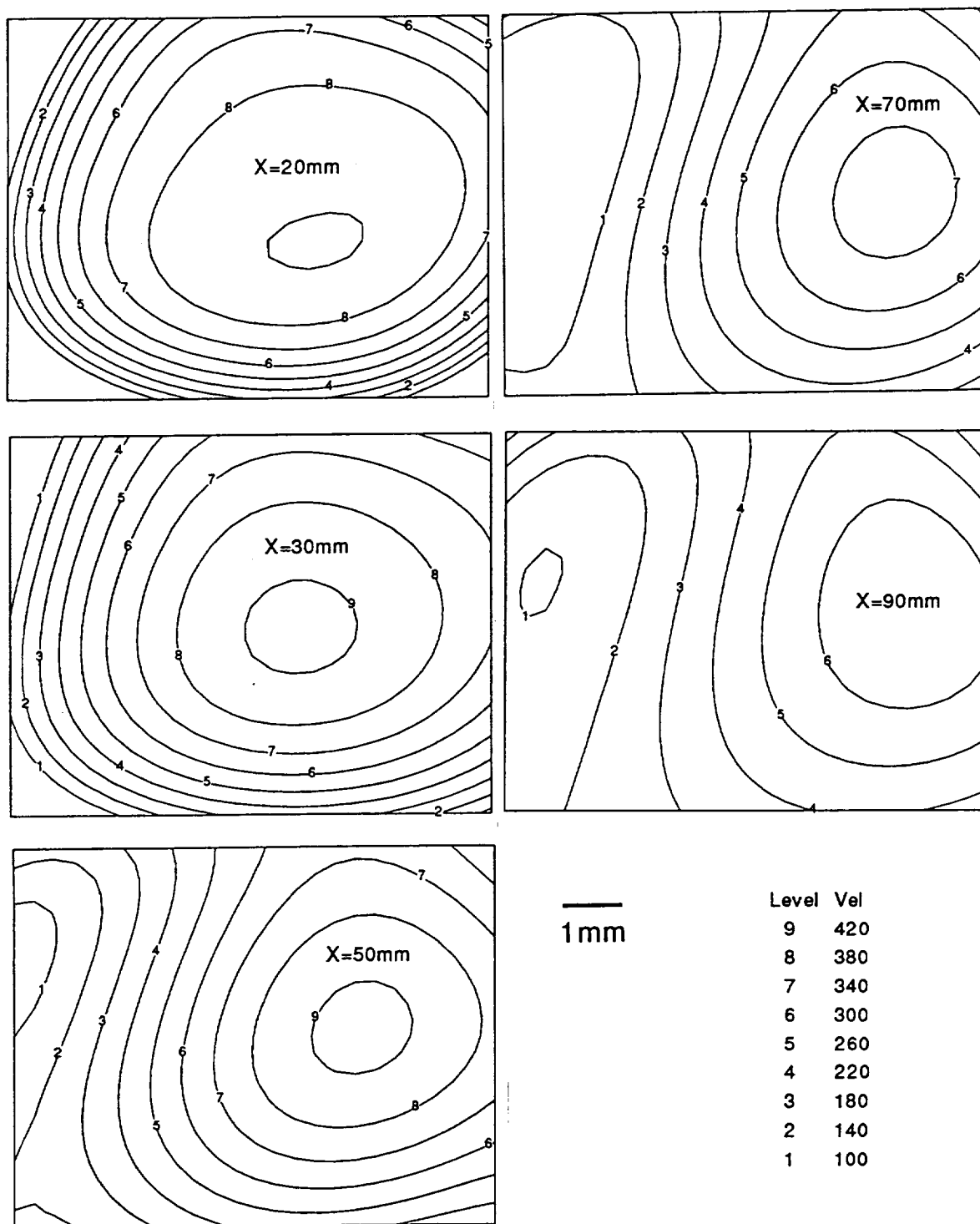


Figure 3. Measured constant axial velocity contours. Distances from nozzle exit are $X=20,30,50,70,90\text{mm}$. Isentropic velocity = 422m/s .

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